

# Expectation maximization method in X-ray solar astronomy

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Starting from count modulation profiles, the reconstruction problem is to find the image  $f$  such that

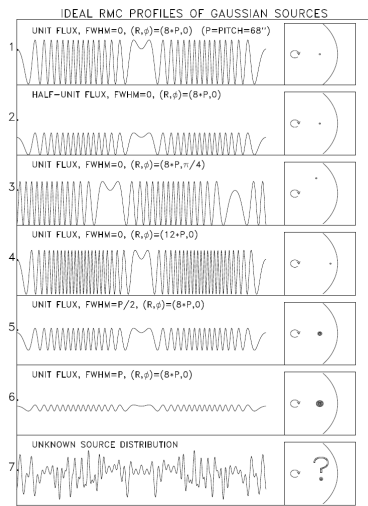
## The model

$$Hf = g$$

where

- $g$  is the detected count modulation profiles.
- $H$  is a linear operator which describes the transformation between images and profiles provided by the RMCs.

# Matrix $H$ mimicks Rhessi production of count modulation profiles



The vector  $g$  can be consider a realization of a Poisson random variable  $Y$  with expected value  $Hf$ .

The probability density function of variable  $Y$

## Poisson probability

$$p(g; f) = \prod_{i \in R} \frac{e^{-(Hf)_i}}{g_i!} (Hf)_i^{g_i}$$

yields the Likelihood  $L(f; g)$  of the problem.

Maximum likelihood method obtains the solution of  $f$  when the probability function reaches its maximum value.

# Constrained minimization

We are interested in non negative solutions.

We minimize the C-statistics ( $\simeq$  negative logarithm of the Likelihood)

## C-statistics

$$C(g, f) = \sum_{i=0}^N g_i \log \frac{g_i}{(Hf)_i} + (Hf)_i - g_i$$

under non negative constraint, i.e.

## Constrained minimization problem

$$\min C(g, f) \quad | \quad f \geq 0$$

# Expectation maximization algorithm

- The necessary and sufficient condition for  $f$  to be the solution of the constrained C-statistics minimization is Karush-Khun-Tucker (KKT) conditions. They lead to a fixed point equation iteratively solved by the EM algorithm

## EM algorithm

$$f^{k+1} = \frac{f^k}{H^T \mathbf{1}} H^T \left( \frac{g}{H f^k} \right)$$

- Property 1 : This algorithm converges to a non negative solution of the problem.
- Property 2 : At each iteration the total number of expected counts is equal to the total detected counts.
- Some regularization is required.

# Regularization specifically conceived for Rhessi imaging

- Classical stopping rules do not efficiently work for Rhessi imaging:

$$r(f, g) \leq \mathbb{E}_Y r(f, g)$$

where  $r(f, g)$  is some approximation of the C-statistics.

- This approach does not efficiently work since it does not take into account the positive constraint on the solution !
- We developed a new rule to regularize EM for Rhessi imaging by taking:

$$r(f, g) = \left\| f \cdot H^T \left( \frac{g}{Hf} - 1 \right) \right\|^2$$



# EM algorithm: applications

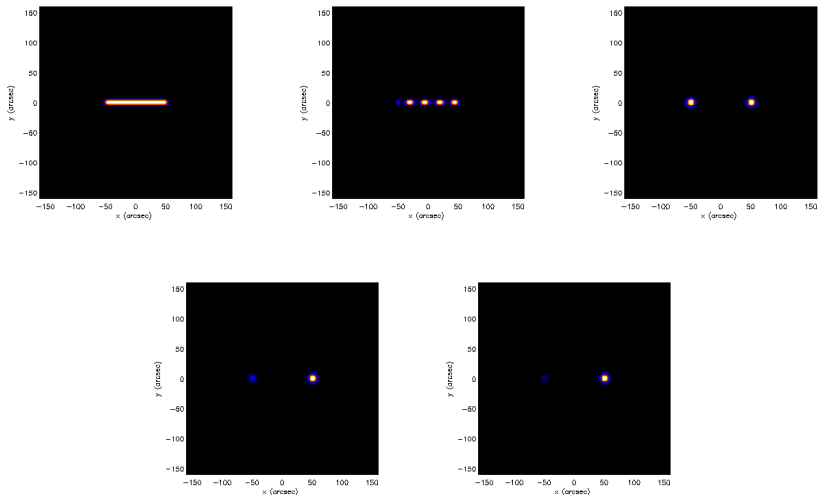


Figure: Annapolis synthetic maps

# EM algorithm: applications

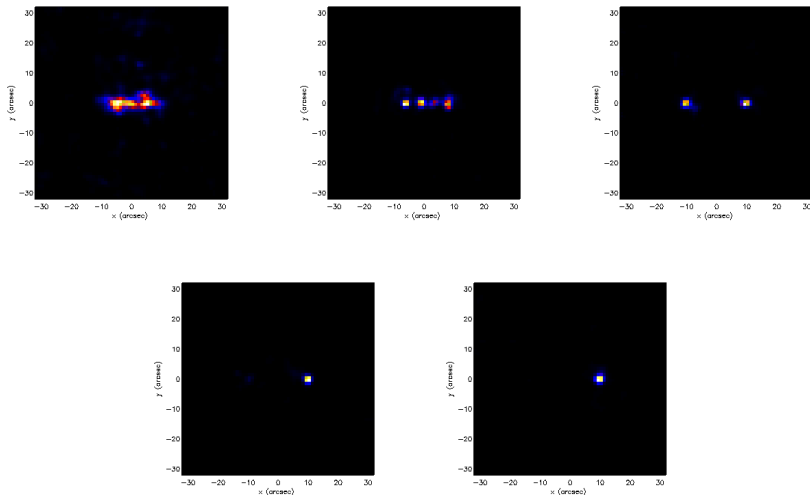


Figure: Low statistics:  $10^3$  counts per detector

# EM algorithm: applications

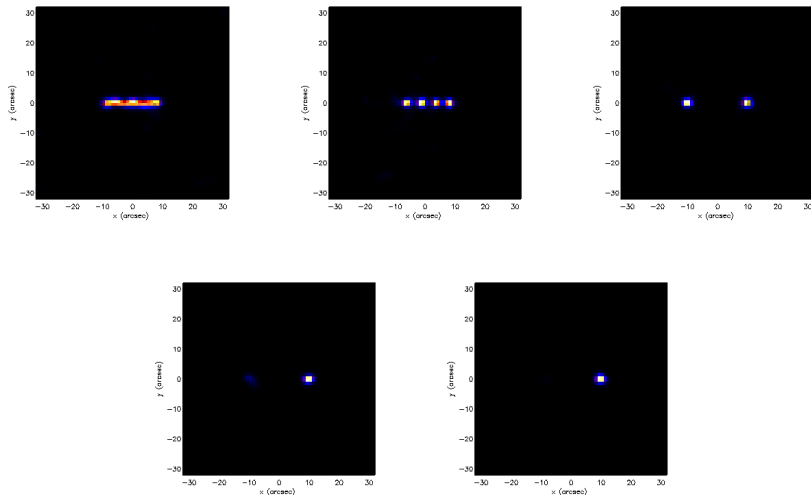


Figure: Medium statistics:  $10^4$  counts per detector

# EM algorithm: applications

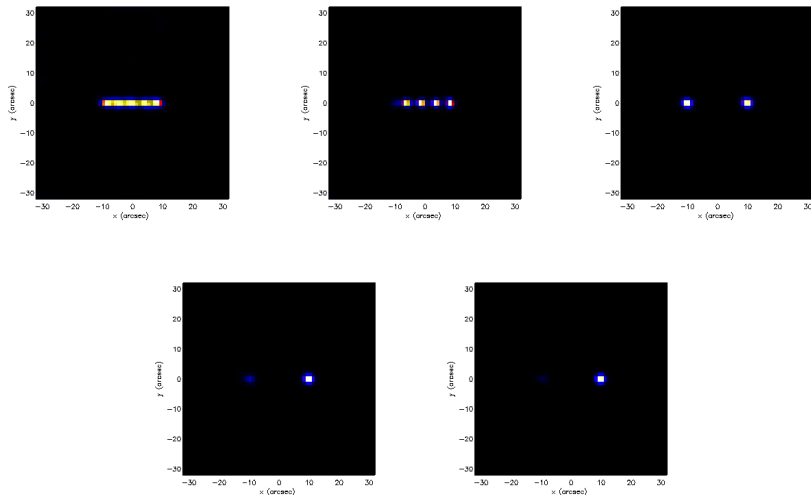


Figure: High statistics:  $10^5$  counts per detector

# EM algorithm: applications

## Numerical comparison with Pixon method:

### LINE SOURCE WITH CONSTANT INTENSITY

#### CRITERIUM A: ORIENTATION (0 deg)

	HIGH	MEDIUM	LOW
ALGORITHM	DET 1-9	DET 1-9	DET 1-9
PIXON	-0.10	-0.61	3.84
EM	-0.06	-0.41	3.33

#### CRITERIUM B: Number of sources at FWHM (1)

	HIGH	MEDIUM	LOW
ALGORITHM	DET 1-9	DET 1-9	DET 1-9
PIXON	1	2	2
EM	1	3	2

#### CRITERIUM C:

##### Number of sources from local minima analysis and $K=0.2$ (1)

	HIGH	MEDIUM	LOW
ALGORITHM	DET 1-9	DET 1-9	DET 1-9
PIXON	1	3	2
EM	1	3	2

### LINE SOURCE WITH VARYING INTENSITY

#### CRITERIUM A: ORIENTATION (0 deg)

	HIGH	MEDIUM	LOW
ALGORITHM	DET 1-9	DET 1-9	DET 1-9
PIXON	-0.10	0.01	0.11
EM	-0.04	-0.06	-0.96

#### CRITERIUM B: Number of sources at FWHM (4)

	HIGH	MEDIUM	LOW
ALGORITHM	DET 1-9	DET 1-9	DET 1-9
PIXON	4	4	3
EM	4	4	3

#### CRITERIUM C:

##### Number of sources from local minima analysis and $K=0.2$ (5)

	HIGH	MEDIUM	LOW
ALGORITHM	DET 1-9	DET 1-9	DET 1-9
PIXON	5	4	4
EM	4	4	4

# EM algorithm: applications

## Numerical comparison with Pixon method:

### LINE SOURCE WITH CONSTANT INTENSITY

#### CRITERIUM D: FWHM (20.2 arcsec)

	HIGH	MEDIUM	LOW
ALGORITHM	DET 1-9	DET 1-9	DET 1-9
PIXON	19.81	18.31	14.47
EM	18,86	17.84	13.51

#### CRITERIUM E: FWHM in the orthogonal direction (1.35 arcsec)

	HIGH	MEDIUM	LOW
ALGORITHM	DET 1-9	DET 1-9	DET 1-9
PIXON	3.04	2.17	4.27
EM	2,08	1.94	3.45

#### CRITERIUM F: rms

	HIGH	MEDIUM	LOW
ALGORITHM	DET 1-9	DET 1-9	DET 1-9
PIXON	0.6082	0.5297	0.7519
EM	0,4865	0.4895	0.6991

### LINE SOURCE WITH VARYING INTENSITY

#### CRITERIUM D: FWHM (16.90 arcsec)

	HIGH	MEDIUM	LOW
ALGORITHM	DET 1-9	DET 1-9	DET 1-9
PIXON	17.82	16.58	17.13
EM	16.41	16.35	15.92

#### CRITERIUM E: FWHM in the orthogonal direction (1.35 arcsec)

	HIGH	MEDIUM	LOW
ALGORITHM	DET 1-9	DET 1-9	DET 1-9
PIXON	3.79	2.81	3.10
EM	2.05	2.17	1.86

#### CRITERIUM F: rms

	HIGH	MEDIUM	LOW
ALGORITHM	DET 1-9	DET 1-9	DET 1-9
PIXON	0.7366	0.6521	0.7574
EM	0.5517	0.5553	0.6655

# EM algorithm: applications

Numerical comparison with Pixion method:

TWO SOURCES: FLUX RATIO 1 - 5 - 10											
FLUX RATIO =1					FLUX RATIO =5				FLUX RATIO =10		
CRITERIUM G1: Distance of the first source from the simulated one (0 arcsec)											
	HIGH	MEDIUM	LOW		HIGH	MEDIUM	LOW		HIGH	MEDIUM	LOW
ALGORITHM	DET 1-9	DET 1-9	DET 1-9		DET 1-9	DET 1-9	DET 1-9		DET 1-9	DET 1-9	DET 1-9
PIXON	0.71	0.71	0.71		1.58	1.58	03.54		1.58	7.52	9.93
EM	0.71	0.71	0.71		0.71	0.71	0.71		0.71	0.71	3.54
CRITERIUM G2: Distance of the second source from the simulated one (arcsec)											
	HIGH	MEDIUM	LOW		HIGH	MEDIUM	LOW		HIGH	MEDIUM	LOW
ALGORITHM	DET 1-9	DET 1-9	DET 1-9		DET 1-9	DET 1-9	DET 1-9		DET 1-9	DET 1-9	DET 1-9
PIXON	0.71	0.71	0.71		1.58	1.58	03.54		1.58	7.52	9.93
EM	0.71	0.71	0.71		0.71	0.71	0.71		0.71	0.71	3.54
CRITERIUM H: separation between the two sources (20 arcsec)											
	HIGH	MEDIUM	LOW		HIGH	MEDIUM	LOW		HIGH	MEDIUM	LOW
ALGORITHM	DET 1-9	DET 1-9	DET 1-9		DET 1-9	DET 1-9	DET 1-9		DET 1-9	DET 1-9	DET 1-9
PIXON	22.02	19.00	19.00		18.00	19.00	24.02		19.00	28.02	27.66
EM	19.03	19.00	20.00		20.00	19.03	20.03		19.00	19.00	23.00

# EM algorithm: applications

## Numerical comparison with Pixon method:

### TWO SOURCES: FLUX RATIO 1 - 5 - 10

#### FLUX RATIO =1

#### FLUX RATIO =5

#### FLUX RATIO =10

#### CRITERIUM I: Orientation of the separation line (0 deg)

ALGORITHM	HIGH	MEDIUM	LOW		HIGH	MEDIUM	LOW		HIGH	MEDIUM	LOW
	DET 1-9	DET 1-9	DET 1-9		DET 1-9	DET 1-9	DET 1-9		DET 1-9	DET 1-9	DET 1-9
PIXON	-2.6	0	0		0	0	-2.39		0	-2.05	-12.53
EM	3.01	0	0		0	-3.01	-2.86		0	0	0

#### CRITERIUM J: Ratio of the two peak intensities (1-5-10)

ALGORITHM	HIGH	MEDIUM	LOW		HIGH	MEDIUM	LOW		HIGH	MEDIUM	LOW
	DET 1-9	DET 1-9	DET 1-9		DET 1-9	DET 1-9	DET 1-9		DET 1-9	DET 1-9	DET 1-9
PIXON	0.95	0.93	1.56		5.77	9.34	121.12		7.72	74.44	471.86
EM	0.94	1.02	1.37		6.79	10.93	21.59		23.82	92.26	245.61

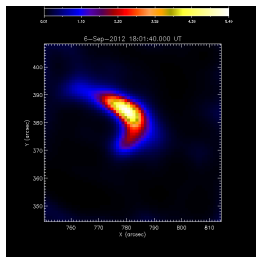
#### CRITERIUM K: Ratio of the two fluxes in the two circles centered in the two sources and with radius =d/2 (1-5-10)

ALGORITHM	HIGH	MEDIUM	LOW		HIGH	MEDIUM	LOW		HIGH	MEDIUM	LOW
	DET 1-9	DET 1-9	DET 1-9		DET 1-9	DET 1-9	DET 1-9		DET 1-9	DET 1-9	DET 1-9
PIXON	1.01	1.02	1.00		5.14	5.57	11.53		11.65	16.96	42.69
EM	1.01	1.01	0.99		4.73	4.99	7.56		10.02	16.15	29.68

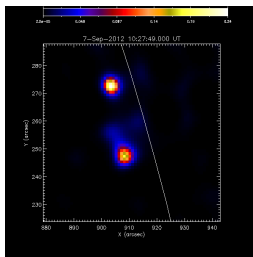


# EM algorithm: applications with real data

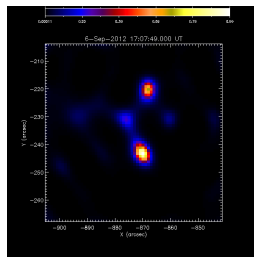
Reconstruction of three different events: a loop, two footpoints and two footpoints with a small loop-top.



(a) Apr 15 2002.  
Time:  
00.06.00 00.08.00.  
Energy:12-14 keV.



(b) Feb 20 2002.  
Time:  
11.05.58 11.06.41.  
Energy:25-30 keV.



(c) Jul 23 2002.  
Time:  
00.30.00 00.32.00.  
Energy:100-300 keV.

# EM algorithm in spectroscopy

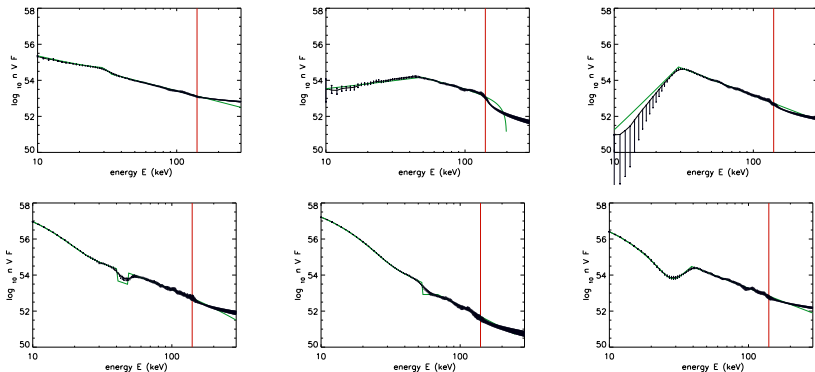
Expectation maximization can be used for spectroscopy.

- $g$  is a given count spectrum.
- $f$  is the unknown electron spectrum.
- $H$  is a matrix containing information about the Bremsstrahlung cross-section and the Detector Response Matrix (DRM).

## Remark

*DRM can take into account different aspects of Rhessi hardware. The inversion results can be different according to DRM choice.*

# Spectroscopy: blind tests



**Figure:** Six different synthetic spectra (green lines) and its reconstruction with EM. The red lines indicates the maximum energy of detected counts.

# EM algorithm for electron flux maps

Our most recent advance is the application of EM for electron maps reconstruction.

- $g$  is a given set of count modulation profiles detected from  $\epsilon_{min}$  and  $\epsilon_{max}$ .
- $f$  is the unknown set of electron maps.
- $H$  is a tensor containing information about energy and space.

## Remark

*In this context, the regularization of EM method provides smooth behaviour in energy and space. The resulting set of electron maps can be spatially integrated computing the total integrated flux spectrum.*

# EM algorithm for electron flux maps

April 15 2002 event: detector 3-8, 64 arcsec images. Energies of the reconstructed electron maps are from 14 up to 90 keV. Count modulation profiles are up to 52 KeV.

Figure: Electron maps

## Conclusion

Expectation maximization is a powerful technique. Starting from count modulation profiles we can reconstruct:

- X-ray photon images.
- X-ray electron spectra.

## Perspective

To reconstruct electron images by means of Expectation Maximization via counts

- X-ray electron images and spectra.

Thank you